

NATIONAL UNIVERSITY OF SCIENCE AND TECHNOLOGY

APPLIED PHYSICS DEPARTMENT

STATISTICAL MECHANICS - SPH 2102

EXAMINATION

BSc HONOURS PART II: DECEMBER 2004

DURATION 3 HOURS

Instructions To Candidates:

1. Answer ALL parts of question 1 in Section A.
2. Answer any THREE questions from Section B
3. Section A carries 40 marks and Section B carries 60 Marks.
4. Show all your steps clearly in any calculation.

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SECTION A

1. (a) Define the following terms:

- |                                |     |
|--------------------------------|-----|
| (i) Microstate and Macrostate, | [3] |
| (ii) Partition function,       | [2] |
| (iii) Intensive variable,      | [1] |
| (iv) Bosons.                   | [1] |

(b) Distinguish between a microcanonical ensemble and canonical ensemble. [4]

(c) Given that the entropy of an ideal gas in the semiclassical limit is given by:

$$S(E, V, N) = Nk \left[ \ln \frac{V}{N} + \frac{3}{2} \ln \left( \frac{mE}{3\pi N \hbar^2} \right) + \frac{5}{2} N \right]$$

where all symbols have their usual meanings, show the following

- (i) Thermal equation of state is given by  $E = \frac{3}{2} NkT$  [2]

- (ii) Pressure equation of state is given by  $P = \frac{NkT}{V}$  [2]
- (d) (i) State any difficulties you are aware of that arise with the use of Maxwell - Boltzmann statistics. [4]
- (ii) State Heisenberg's principle as it relates to quantum mechanics of an electron gas. [2]

- (e) Given that the Maxwell - Distribution for molecular speeds is given by the

$$\text{expression } f(v) = 4\pi n \left( \frac{m}{2\pi kT} \right)^{3/2} v^2 e^{-mv^2/2kT}$$

Use the change - of - variable techniques to convert the above expression into an energy distribution function. [4]

- (f) According to Pierre Curie's theory of Paramagnetism, magnetisation is given by the expression  $M = C \frac{H}{T}$ , where the other symbols have their usual meanings. What could be one weakness of this formula? [2]

- (g) (i) Explain why the entropy of a system is sometimes described as a measure of chaos or disorder of that system. [4]

- (ii) Show that the thermodynamic probability  $(\Omega) = \frac{N!}{\pi_i N_i!}$  can be used to derive the expression  $\ln(\Omega) = N \ln N - \sum N_i \ln N_i$  with the use of Stirling's approximation. [4]

- (h) In the vicinity of the triple point, the liquid-vapour coexistence curve of liquid ammonia can be presented by  $\ln P = 24.38 - 3063/T$ , where the pressure is given in Pascals. The vapour pressure of solid ammonia is  $\ln P = 27.92 - 3754/T$ .

- (i) What are the temperature and pressure at the triple point? [2]
- (ii) Calculate the enthalpy of vaporisation. [2]

### SECTION B

2. (a) Consider a system of  $N$  particles in a phase space of 3 cells 1,2 and 3. Suppose that the energy of a particle in cell  $i$  is given by:  
 $W_1 = 0$ ,  $W_2 = W$  and  $W_3 = 2W$ .
- (i) Write down the expressions for the partition function ( $z$ ) and number of phase points in the 3 cells in terms of the characteristic temperature  $\theta$ . [5]
- (ii) Find an expression for the total internal energy [4]
- (iii) Find an expression for the entropy [4]
- (b) Describe the distribution of the particles under the following conditions.
- (i) At temperatures that are very small compared to the characteristic temperature [4]
- (ii) At temperatures that are very large compared to the characteristic temperature. [3]
- 3 (a) (i) Show that the Langevin theory for paramagnetism  
$$\left[ M = n\mu \left( \coth \frac{\mu B}{kT} - \frac{kT}{\mu B} \right) \right]$$
lead to Curies law in weak fields and at high temperatures.  
(Given:  $\coth x \approx \frac{1}{x} + \frac{x}{3}$  for a small angle) [6]
- (ii) Write down the expression for the resulting curie's constant  $C$ . [2]
- (iii) Explain why Langevin's theory can be said to be a more appropriate approach compared to other approaches for Paramagnetism. [3]
- (b) Describe the Stern-Gerlach experiment and explain how it is used to demonstrate

space quantization. [9]

4. (a) Describe briefly the Einstein model of a solid. [3]

(b) Given that the partition function of a single quantum harmonic oscillator is given

by: 
$$Z = \frac{\exp(-\beta\hbar\omega/2)}{1 - \exp(-\beta\hbar\omega)}$$

For the single oscillator determine the expressions for :

(i) The Helmholtz thermodynamic potential [3]

(ii) Entropy [3]

(iii) Mean energy [3]

(iv) Hence determine the heat capacity of an Einstein solid. [3]

(v) Determine the limiting behaviour of the heat capacity in (iv) at low temperatures and at high temperatures and comment on its consistency with the observed behaviour of solids. [5]

5. (a) Compare and contrast classical and quantum statistics. [4]

(b) (i) By finding the thermodynamic probability of 3 phase points a, b, and c in 2 cells, i and j both under Maxwell - Boltzmann statistics and under Bose - Einstein statistics.

Show the difference between the two if the Macro state is  $N_i = 2$ ,  $N_j = 1$ .

You should use diagrammatic illustrations. [6]

(ii) Use the following 2 assumptions

(1) Compartments in cell i are numbered 1, 2, 3, -----n.

(2) Phase points are lettered a, b, c, ----- $N_i$

to explain why the expression for thermodynamic probability is given by:

$$(\Omega) = \pi \left( \frac{n + N_i - 1}{(n - 1)! N_i!} \right)! \text{ for Bose - Einstein statistics. [5]}$$

- (iii) Use Stirling's approximation and Lagrange's method of undetermined multipliers to show that the expression for Bose - Einstein distribution function is given by  $\frac{N}{n} = \frac{1}{B \exp(\beta \epsilon_i) - 1}$  where the all symbols have their usual meanings. [5]

6. (a) For a black body radiation, explain the following:

- (i) Spectral energy density  $u(E)dE$  [3]  
 (ii) Stefan's Law [2]  
 (iii) Wien's displacement law [2]

(b) (i) Statistical mechanics shows that at temperature T, the distribution function

for a photon gas is  $\frac{N}{n} = \frac{1}{\exp(\epsilon_i / kT) - 1}$

where all symbols have their usual meanings. Use this fact to show that the energy density within a frequency range  $\nu$  is

$$\frac{h\nu dN\nu}{V} = \frac{8\pi h\nu^3}{C^3} \cdot \frac{1}{\exp(h\nu/kT) - 1} d\nu : \text{Planck's law.} \quad [8]$$

- (ii) Show how the above Planck's formula can be reduced to Rayleigh - Jeans's classical law for black body radiation at long wavelengths.

[5]

- END OF EXAMINATION -